



Pegasus-III
Experiment

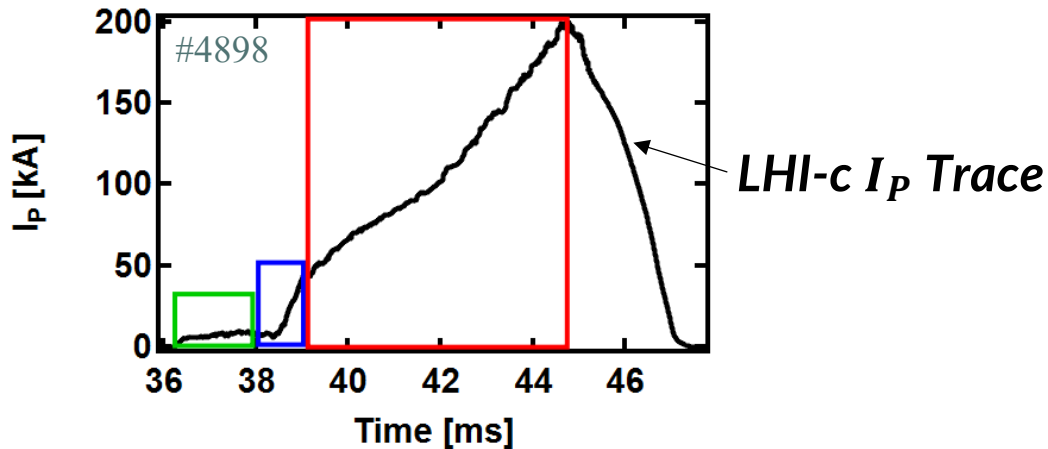
Development of a Microwave System for Pegasus-III

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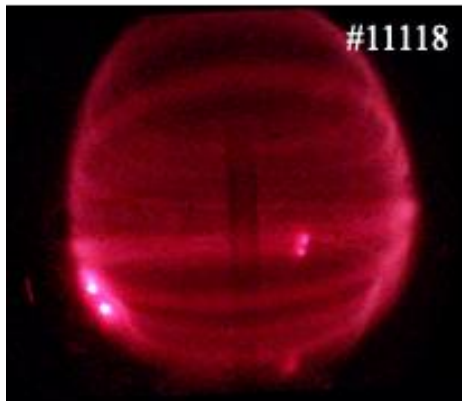


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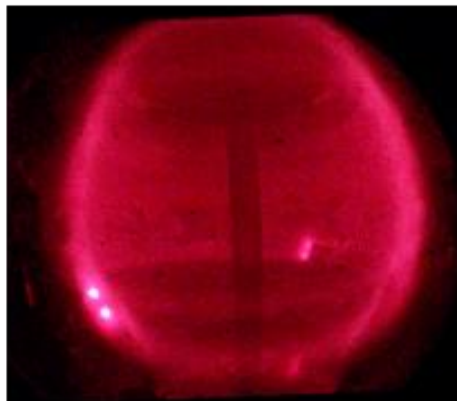
The Pegasus-III Experiment



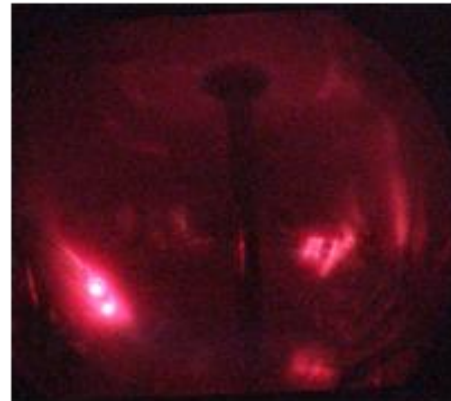
Fast Camera Images of LHI Stages on Pegasus-III



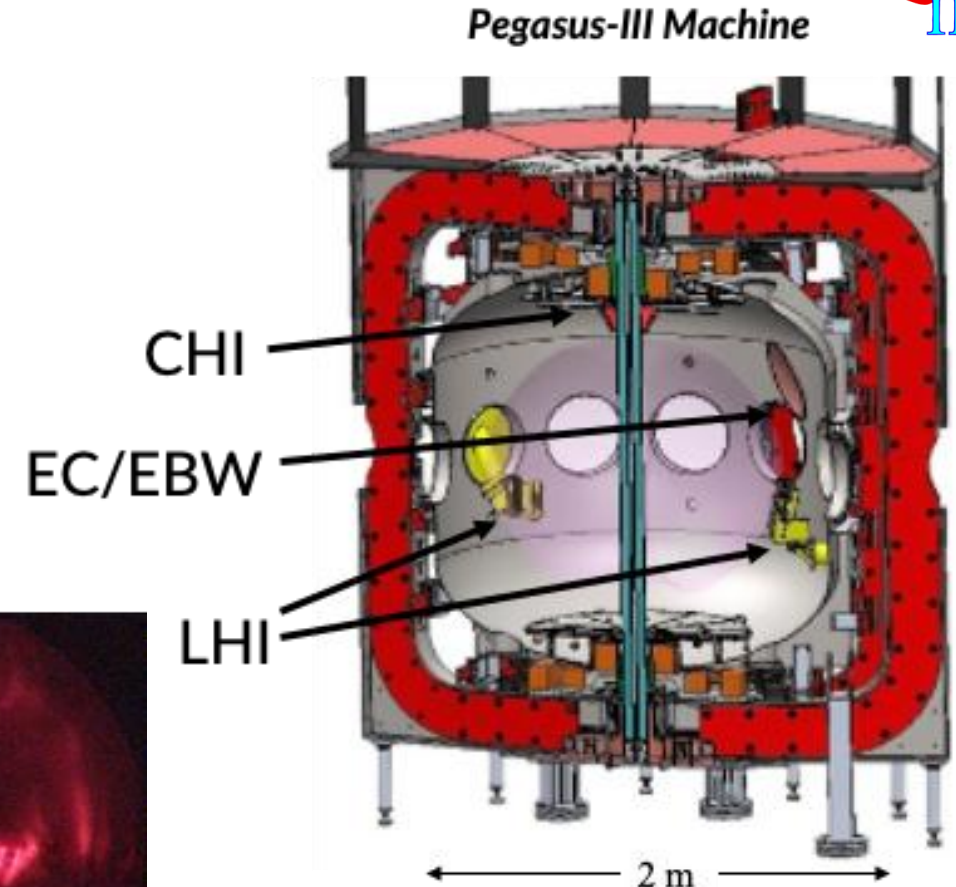
$I_p \approx N_{turns} I_{inj}$
1. Injected current follows helical \vec{B}



$I_p \gtrsim N_{turns} I_{inj}$
2. Current streams go unstable, reconnect



$I_p \gg N_{turns} I_{inj}$
3. Plasma relaxes to a "tokamak-like" state



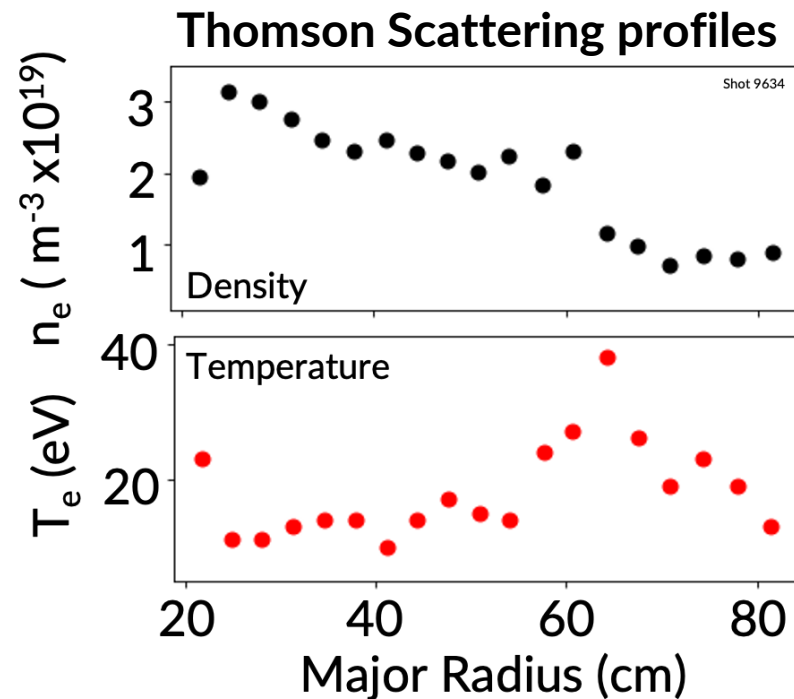
Typical parameters:

- $n_{e0} \sim 2 \times 10^{19} \text{ m}^{-3}$
- $T_e \sim 30 - 100 \text{ eV}$

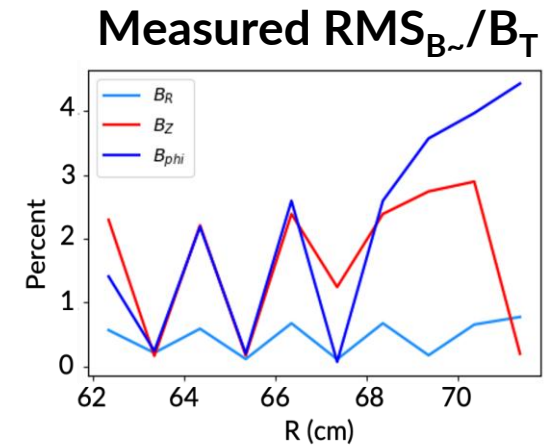
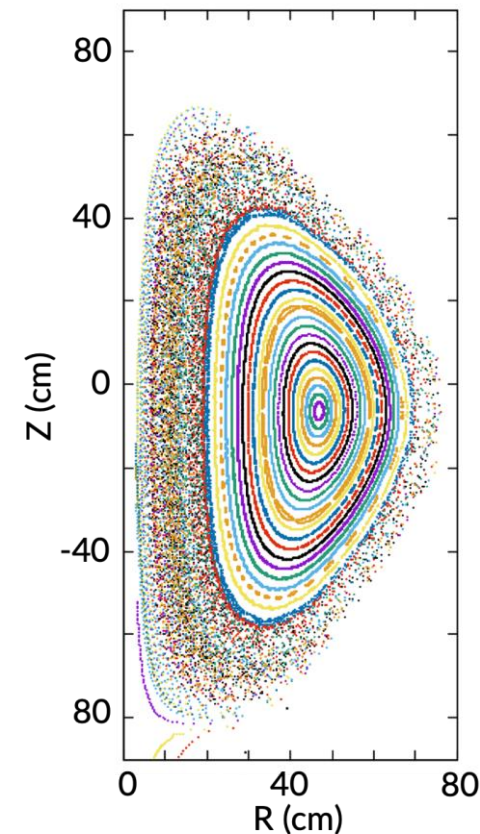


Characteristics of LHI plasmas

- Hollow electron temperature profiles
- Peaked or flat electron density profiles
- Stochastic edges, B-field fluctuations $\sim 3\%$



Poincaré plot calculated with M3D-C1¹



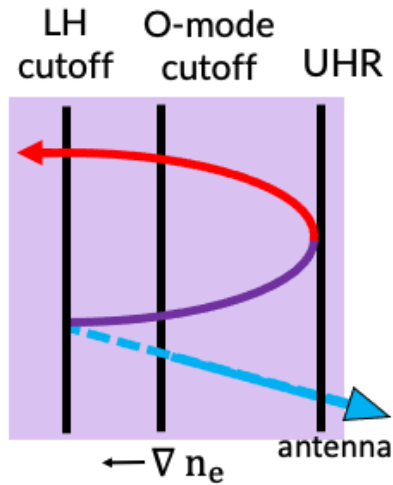
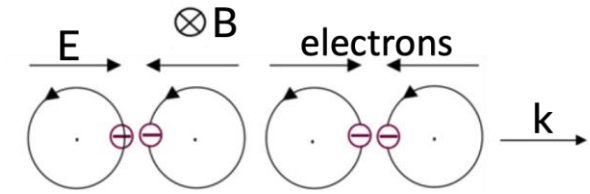
¹C.E. Schaefer, et al, 2025. [Online].
<https://arxiv.org/abs/2511.03930>



Electron Bernstein Waves

- For a 28 GHz source, plasmas become overdense at $1 \times 10^{19} \text{ m}^{-3}$, prohibiting EC use. Instead, EBW can be employed

- Do not experience n_e cutoffs, are absorbed at EC harmonics



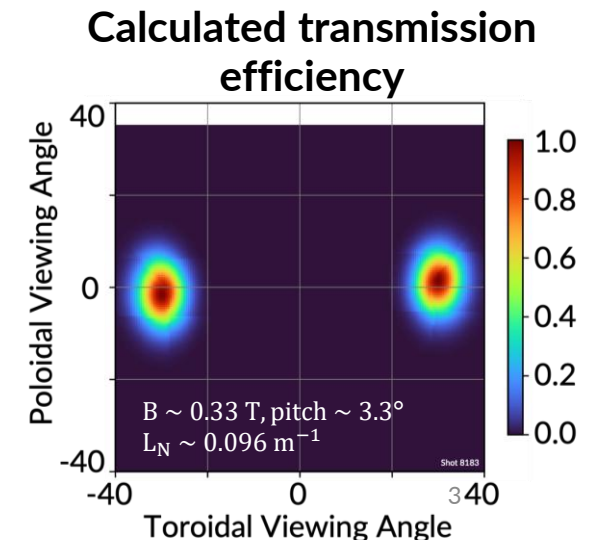
- EBW cannot propagate in vacuum and are launched via OXB mode conversion (MC) scheme

- Launched **O-mode** converts to **X-mode** at LH cutoff, **X-mode** to **EBW** conversion occurs at UHR layer

- Power coupling depends on B field and $L_n = \frac{n_e}{dn_e/dr}$

- When $T_e > 1 \text{ eV}$ and $n_e > 10^{11} \text{ m}^{-3}$ plasma acts as blackbody EBW emitter.

- Through BXO MC, EBW are emitted
- OX conversion occurs at two angular locations in plasma edge
- Size & angular location of windows estimated with analytical eq.





EBW emission measurements

- Measurements on Pegasus-III have been taken with the SAMI phased-array radiometer
 - SAMI has measured microwave power from a variety of plasmas
 - No MC windows have been conclusively identified yet

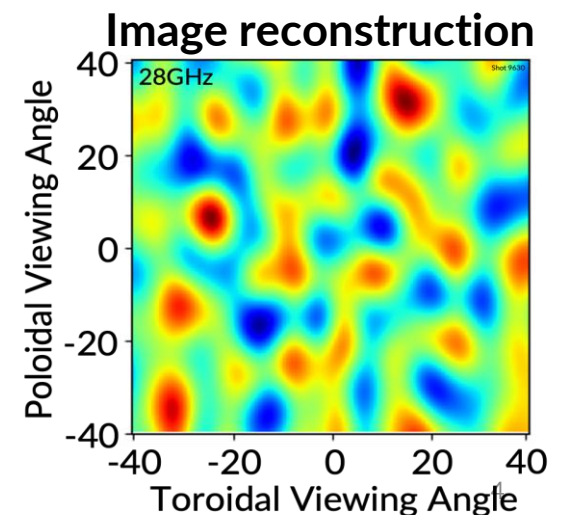
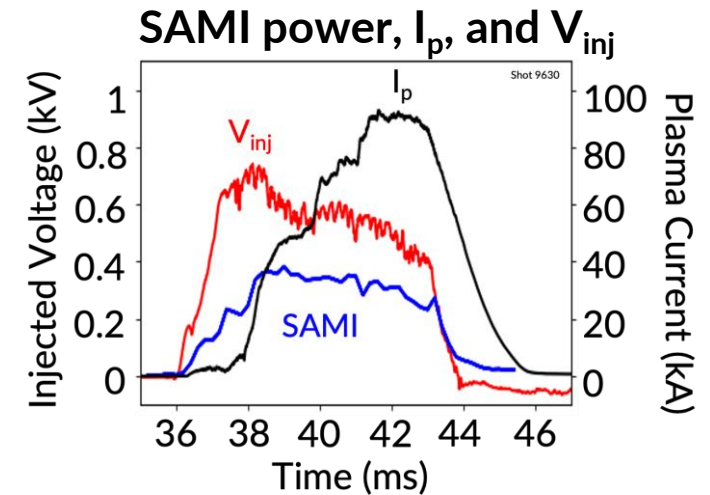
Optical thickness and expected power emission from streams (~800 eV)

| ω (GHz) | n_{harm} | τ_X | X_{Power} (mW) | τ_0 | O_{Power} (mW) | τ_{EBW} | $\text{EBW}_{\text{Power}}$ (mW) |
|-------------------|-------------------|-----------------|----------------------------|------------------|----------------------------|---------------------|-------------------------------------|
| 18 | 4th | $2.3\text{e-}3$ | $3.0\text{e-}3$ | - | 0 | 2768.1 | 1.3 |
| 28 | 6th | $5.3\text{e-}5$ | $2.6\text{e-}4$ | $1.2\text{e-}12$ | 0 | 2715.9 | 5.0 |

And from bulk plasma (~30 eV)

| | | | | | | | |
|----|-----|------------------|------------------|---|---|--------|-----------------|
| 18 | 4th | $1.2\text{e-}8$ | $3.0\text{e-}8$ | - | 0 | 1455.1 | $1.7\text{e-}2$ |
| 28 | 6th | $1.3\text{e-}12$ | $1.2\text{e-}11$ | - | 0 | 1427.6 | $6.4\text{e-}2$ |

- Optical thickness calculations say emission is EBW, so why are there no MC windows?

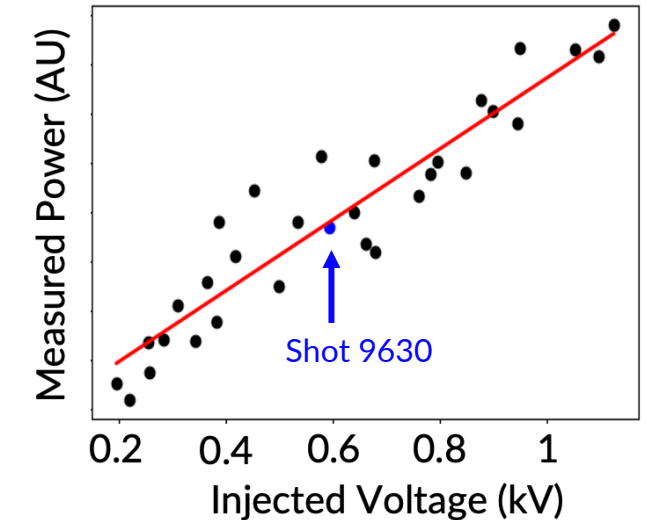




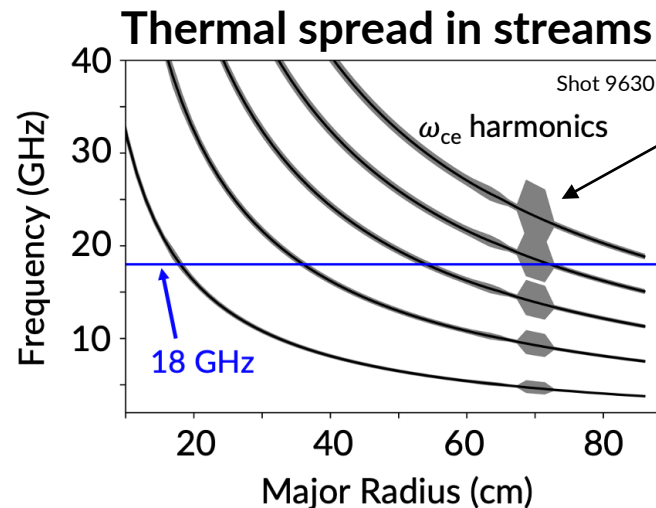
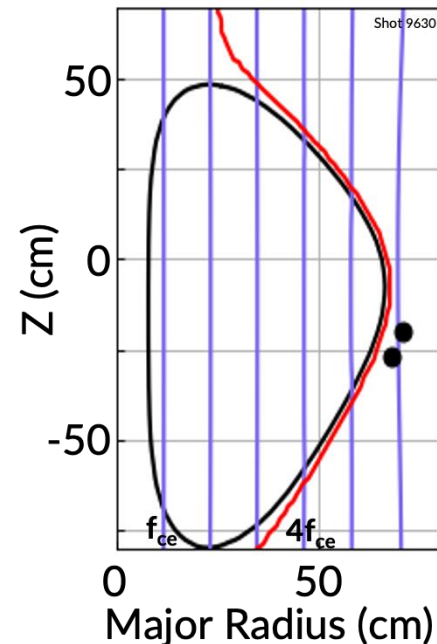
EBW emission measurements from LHI streams

- Scans of injector voltage show measured signal levels are proportional to V_{inj} , suggesting emission from streams dominates measurements
 - EBE is blackbody and directly proportional to T_e for $\tau > 1$
 - T_e in streams exiting injectors is expected to be directly proportional to $V_{inj} \rightarrow T_e[\text{eV}] = V_{inj}[\text{V}]$

Measured 18 GHz power $\propto V_{inj}$



LCFS, MC layer & Harmonics



Streams are
~800 eV

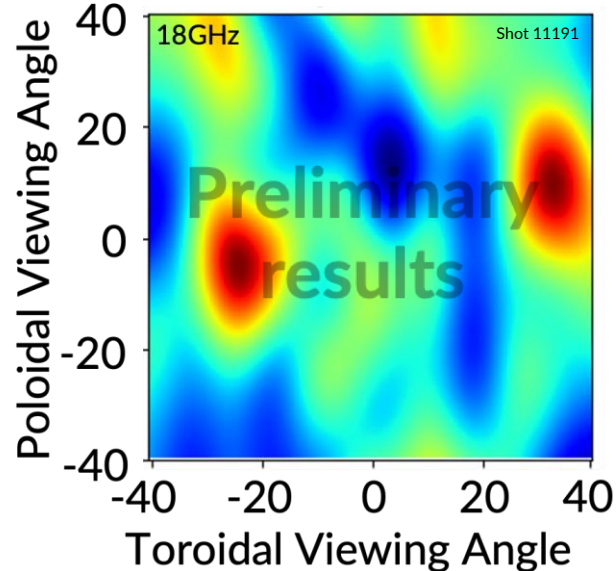
- Thermal spread $\omega_{ce} / (1 - 3N_{\parallel} \frac{v_{th}}{c})$ of streams allows intersection with many harmonics
- Streams are 100x brighter than bulk plasma, limiting search for MC windows to the decay phase



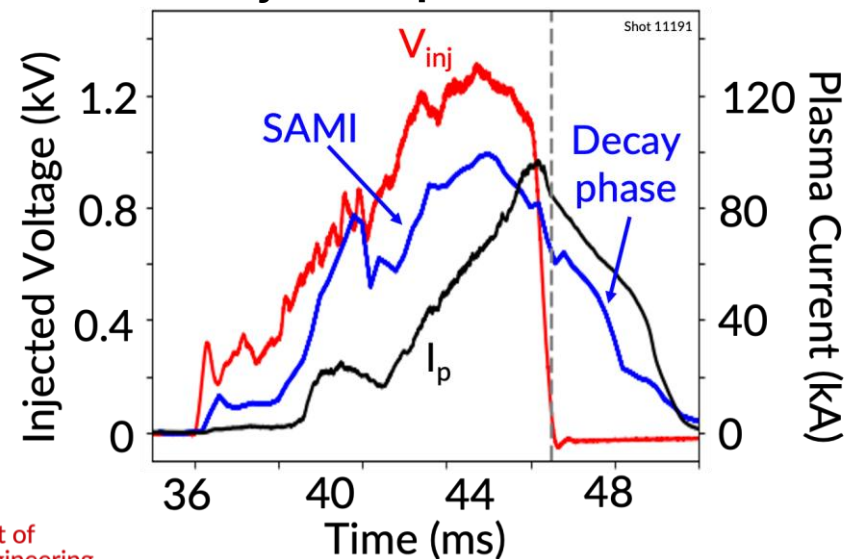
Promising scenario for MC window measurement

- MC windows were potentially seen in SAMI measurements during decay phase
- **NO Thomson Scattering measurements of these shots** -> cannot confirm MC layer or window locations. Results are preliminary only
- What's different about this scenario:
 - Single injector + high V_{inj} -> very low fueling
 - Low fueling -> higher T_e , lower edge neutral density & collisional damping

Image reconstruction at 44.7 ms



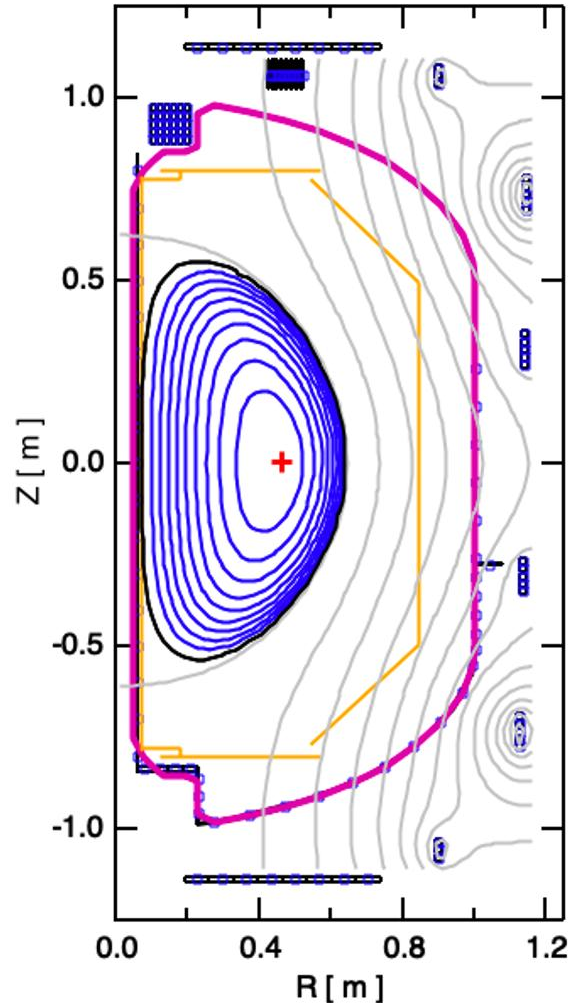
1-inj SAMI power trace



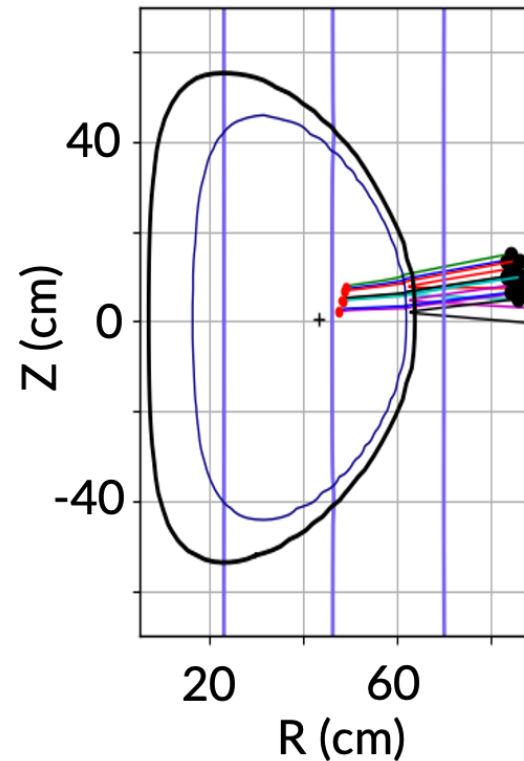
Predictive 0.6 T equilibrium and representative 0.3 T Thomson Scattering profiles



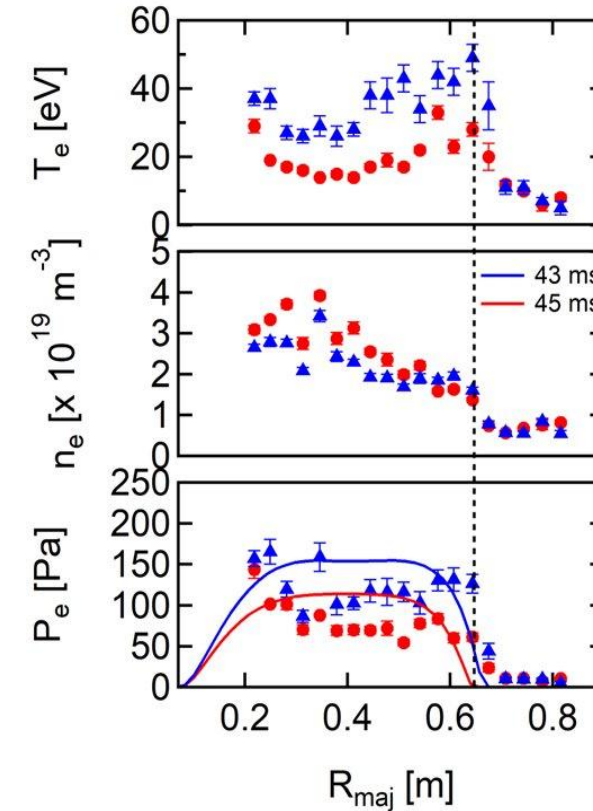
Predictive 0.6 T Equilibrium



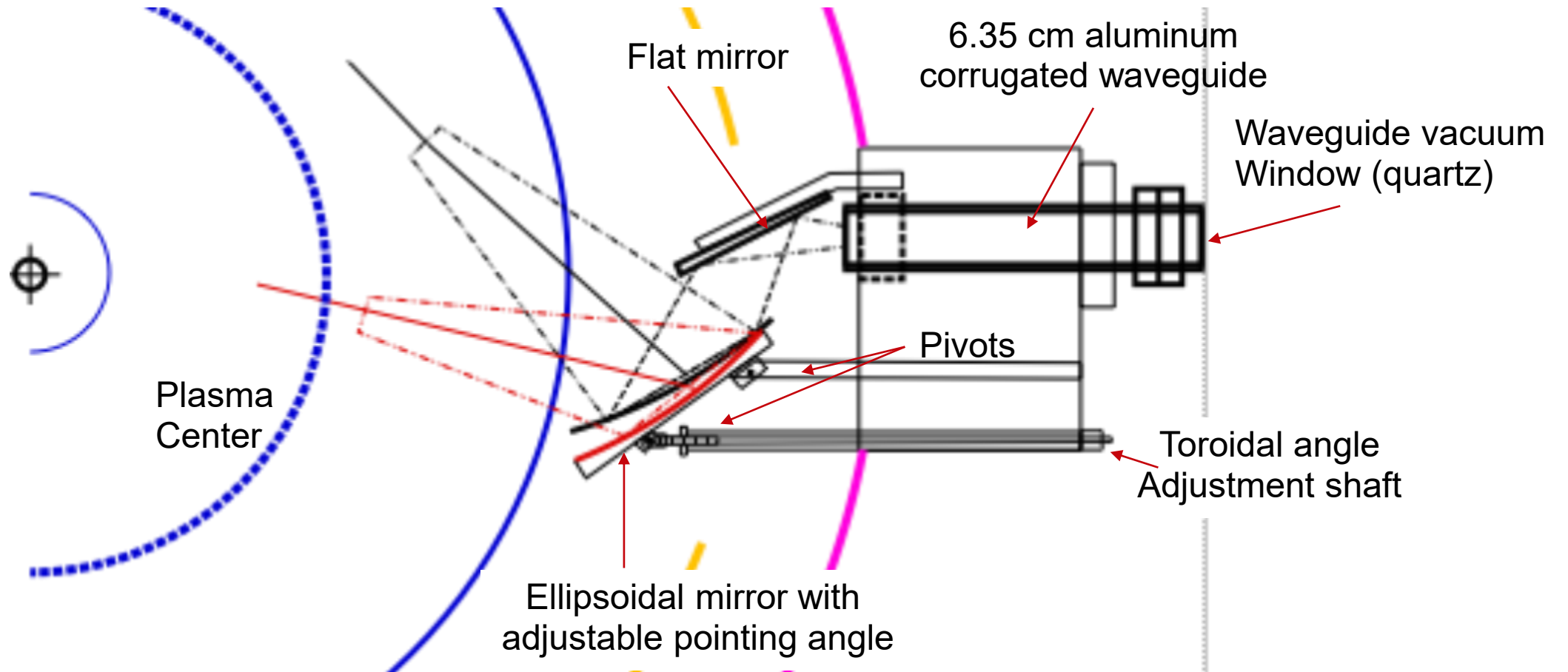
28 GHz harmonic locations



TS Profiles for 0.3 T Shot



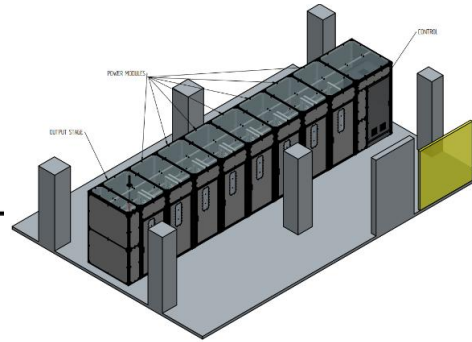
Proposed launch system





Proposed equipment layout

- Launch possible from midplane or top ports
- ~ 15 m across the hall
- 85kV power supply for 150 kW, 28 GHz gyrotron



Gyrotron
Power Supply

Chillers

Magnet
Supplies

28 GHz
Gyrotron

480
Panel
225A
6 slot

208 30A
10kW
Chiller

480 50A
60kW
Chiller

Room
B154



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