

How To Make a Blue Plasma: Impurities in WHAM

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Tiny Plasma + Lots of Heating = Impurities





- Identifying impurity elements
 - Optical spectroscopy (visible range)
 - Collisional radiative modeling
 - Residual gas analysis
 - Soft X-Ray pulse height spectroscopy
- Identifying impurity sources
 - Inspection of plasma facing components (PFCs)
 - High speed and long exposure videos of plasma
- Future work

Optical Emission Spectroscopy

HIN Optical Emission Spectroscopy at Mid-Plane

- ORNL Portable Diagnostics Package (PDP) [1]
- Mainly for Thomson scattering
- 1x fiber bundle with 10 sightlines dedicated to Optical Emission Spectroscopy (OES)
- 2x Teledyne Princeton Instruments IsoPlane 320 Czerny-Turner spectrometers
 - Instrument function with 2400 I/mm grating ~0.1 nm
 - Pixel pitch ~0.008 nm
 - Span ~ 8 nm
 - Adjustable gating and amplification with ICCD
- One exposure per shot
 - Repeat many shots to get rough time-dependence
- Visible only: limited by window and fiber ~400-800 nm
 - ~30 shots to scan across entire visible range
- 2x Ocean Optics USB spectrometers

[1] K. Fujii et al. "Application of a Portable Diagnostic Package to the Wisconsin HTS Axisymmetric Mirror" Rev. Sci. Instr. (under review)





A Very Active Visible Spectrum!

- H-Balmer series by far the brightest
- Very bright line from C-III at 465 nm
- \rightarrow Large dynamic range required
- \rightarrow Resolution >0.5 nm is insufficient to accurately identify impurity peaks









Initial Identification: Fluorine and Carbon

- C-III line always present and very bright with ECH heating
- Strong fast electron tail with low pressure \rightarrow F-I line appears
- → Removal of Teflon components near the plasma reduced F-I brightness to near zero
 - Flux loop support and edge electrostatic probes
- \rightarrow Removal of 3D printed plastic reduced C-III brightness, but not to zero
- \rightarrow X-rays from fast electrons hitting sensor directly cause sharp lines









After Baking and Conditioning: Still, Lots of Lines

- Concentrated in the UV/blue range for high-Z impurities: Fe, O, Cu, N tentatively identified
 - Wavelength calibration of spectrometer is non-linear
- · Shot-to-shot repeatability is good for plasma parameters, but not necessarily for spectra
 - X-rays? UFOs? Wildly different dynamic range required for different regions in visible
- \rightarrow Collisional Radiative Model required for accurate interpretation





CR Model in Thermal Equilibrium Is Insufficient

- WHAM should be hot enough to burn through C-III
 - Assuming thermal equilibrium with constant population (zero source, zero sink)
- But only 20 ms pulse length + heating power correlates with both source rate and temperature
- ColRadPy used to model the thermal equilibrium case
- → Even at low temperatures, radiated power can be extreme!





- High-Z impurity lines in the UV and soft x-ray range is invisible to OES spectrometer
- Use of AXUV diodes and soft x-ray pulse height spectroscopy as proxies for impurity level
 - Current filter in 40-70 nm
- ECH heating increases core AXUV emission
 - Hot electron tail or bulk electron temperature increase?
- Low-Z impurities lines always there



AXUV Signal and Radial Emissivity Profile with ECH+NBI and NBI only



Predicted W Spectrum as Function of Temperature in Thermal Equilibrium



- \rightarrow Source rate of gas and impurities depends on heating power
- \rightarrow More heating may not be able "burn through" impurities if sources are not under control
- Adding more gas puff tends to suppress C-III line, likely by lowering bulk temperature





- Doppler broadening and shifting allows determination of impurity ion temperature ~30-50 eV
- Radial bias driven rotation up to ~10-20 km/s
- Phantom video shows plasma immediately turns blue upon NBI turn-on
- \rightarrow Possible beam dump recycling? Effect of biasing?



Shot 250317078



Residual Gas Analysis



RGA Cannot Identify Non-Gaseous Compounds

- SRS-100 RGA logged at 3 Hz frequency is sufficient to identify some gaseous impurities
- Requires significant distance and magnetic shielding to obtain accurate partial pressures
- $\rightarrow H_2O, N_2 \text{ or } CO, O_2, CO_2, CH_4$ are tracked and logged 24/7
 - No baking for this experiment, but carbon compounds are always observed with/without baking
- \rightarrow Suitable for tracking long term trends and carbon content between shots



Soft X-ray Spectroscopy

Pulse Counting Detector Resolves ~1 keV X-rays

70 mm² FastSDD*

C2 (SiaNa) Windo

- Amptek Silicon Drift Detector SDD-123 acquired second-hand
 - 12 um Beryllium window, with fast count rate option to 1 MCPC
 - C2 window is transparent to XUV photons and can overwhelm detector
- Magnetic shielding allowed electronics to function
 - Installed at NW 4-5/8" port looking at ECH launcher
- Configured for low gain, max. count rate
 - Max cooling, achieving -43 C in field
 - 8192 Multichannel Analyzer bins
 - Fastest shaping time: 0.4 us
- Start integration at t=-20 s, integrate for 30 seconds
- Careful calibration and tuning require
 - In-situ due to background magnetic field
 - Sensitive to baking





Confirmed Iron lines 6.4 keV, 7.06 keV

SXR Spectrum Averaged over shots 250319061-250319079



Measurement of Hot Electron Tail/EEDF

- \rightarrow Clear distinction between bulk and tail population: two slopes
- Bulk T_e still too low to be an effective bulk electron temperature measurement
 - At least will require calibration against Thomson scattering
- Line-of-sight is limited, does not see any high-Z PFCs directly
- Inversion of EEDF in-progress
- Observation of **Mo** K_{α} **lines**
 - Disappeared/weakened after limiter fix





Identifying Impurity Sources

Post-mortem of PFCs

Diagnostics-Rich Plasma







Unintended PFC/Limiting Surfaces







Metal PFCs Could be Source of Carbon?





Cu 2p, Fe 2p /2p3, Cr 2p, Ru 3d, Sn 3d5, Zn

Videos of UFOs



Long Exposure Video Reveal UFOs

- \rightarrow Cheap but highly integrated USB webcams can function at extremely high magnetic fields > 7 T
- Streaks never goes in front of foreground objects
- Comes from all directions
- Some curve in the magnetic field



Large Amount of Dust and Debris Recovered









Future Work



- Time-resolved spectroscopy data
 - Use of filterscopes to digitize intensity of strong lines of impurities (C, Fe, N, O)
- High resolution, time-resolved spectrometer dedicated to OES (?)
 - McPherson 1.33 m Czerny-Turner spectrometer with 5 pm resolution
 - iStar 334T intensified CCD to give 4 pm pixel resolution and 55250 spectra per second
- Line ratio measurement for edge temperature estimation and impurity transport
- Improved IR and visible imaging of all PFCs
 - FLIR Boson camera and Basler CMOS camera modules
- Time-resolved soft x-ray data
- Installation of multiple WISP gauges











Future PFC Strategy and Tracking Effects

- Removal of all plastics
- Thorough cleaning and inspection of metal surfaces
- Systematic removal of PFCs:
 - Probe array: reduced fluorine lines to nearly zero, and reduced carbon line brightness
 - Gas baffle: further reduction in carbon line brightness
- Full Ti wall gettering
- Beam dump improvements
- Conditioning with glow discharge and Ion Cyclotron Wall Conditioning (ICWC)
- Remove high-Z impurities in favor of carbon/low-Z materials?
 - Graphite limiter? BN protection?
- Use 316LN wherever possible
- Time-dependent CR model to aid analysis of OES spectrum

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